- Local Partitioning of the Overturning Circulation in
- <sup>2</sup> the Tropics and the Connection to the Hadley and

# <sup>3</sup> Walker Circulation

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Abstract. Conceptually, it is useful to partition the three-dimensional 4 tropical circulation into meridional and zonal components, namely the Hadley 5 and Walker circulations. The averaging involved in their definitions can in-6 troduce ambiguities. These problems can be circumvented by first partition-7 ing the total vertical mass flux into components associated with overturn-8 ing in the meridional and zonal directions respectively, called here the local 9 Hadley and local Walker circulations. Defining the local Hadley and local 10 Walker circulations this way ensures the pair of two-dimensional overturn-11 ing circulations can be added to give the original three-dimensional circu-12 lation, even when the averages are taken over limited domains. The method 13 is applied to the vertical motion from the ERA-Interim reanalysis for the pe-14 riod 1979 to 2009. One important result is that the local Hadley circulation 15 responds much more strongly to ENSO than the local Walker circulation, 16 even though the local Walker circulation in the central Pacific weakens dur-17 ing El Niño years and strengthens and widens during La Niña years. 18

## 1. Introduction

Conceptually, it has proven to be very useful to partition the tropical atmosphere into 10 two independent orthogonal overturning circulations. The first of these, the Hadley cir-20 culation, is commonly defined as the zonally-averaged circulation [e.g. Hartmann, 1994]. 21 The second, the Walker circulation, represents the zonal asymmetries in the tropical circu-22 lation. The Walker circulation is often, but not always, defined by the meridional average 23 of the tropical circulation [e.g. Hartmann, 1994], and its definition is often restricted to 24 the tropical Pacific region [e.g. Julian and Chervin, 1978; Vecchi et al., 2006; Power and 25 Smith, 2007; Tokinaga et al., 2012; L'Heureux et al., 2013]. 26

The Hadley circulation is thought of as the meridional response of the wind field to 27 zonally-averaged net heating in the tropics. It plays a central role in our understanding of 28 the energy and water budgets of the Earth [e.g. Barry and Carleton, 2001; Trenberth and 29 Stepaniak, 2003]. Likewise, the Walker circulation is thought of as the zonal response of 30 the wind field to zonal asymmetries in the meridionally-averaged net tropical heating. It 31 is thought to be inextricably connected with the El Niño Southern Oscillation (ENSO), 32 a connection often expressed through the Southern Oscillation Index (SOI). Both the 33 Hadley and Walker circulations are closely connected to the Asian and Australian mon-34 soons. The huge importance of the Hadley and Walker circulations in our description and 35 understanding of weather and climate have motivated numerous studies on their dynam-36 ics [e.g. Held and Hou, 1980; Gill, 1980; James, 1994], their variability on interannual 37 or interdecadal timescales [e.g. Oort and Yienger, 1996; Chen et al., 2002; Vecchi et al., 38 2006; Yu and Zwiers, 2010], and trends in their position and strength [e.g. Quan et al., 30

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40 2004; Mitas and Clement, 2005; Zhao and Moore, 2008; Hu and Fu, 2007; Stachnik and

<sup>41</sup> Schumacher, 2011], including possible changes under enhanced greenhouse conditions [e.g.

<sup>42</sup> Lu et al., 2007; Seidel et al., 2008; Gastineau et al., 2009; Kang and Lu, 2012].

Although partitioning the tropical atmosphere into two orthogonal overturning circula-43 tions is qualitatively simple, it is more difficult quantitatively [Tanaka et al., 2004; Zhao 44 and Moore, 2008]. The main problem is that, in general, the two orthogonal overturning 45 circulations are interconnected, as variations in the horizontal velocity in both the zonal 46 and meridional directions contribute to the horizontal divergence. Consequently, part of 47 the vertical motion is attributable to the circulation in zonal direction and part to that in 48 the meridional direction. Unless the vertical motion is partitioned carefully, the orthogonal 49 two-dimensional circulations need not sum to the original three-dimensional circulation. 50 Resolving this ambiguity is especially important when assessing changes to the Hadley 51 and Walker circulations. Thus, the main aim of the present paper is to provide a more 52 rigorous basis for the common and useful practice of partitioning the three-dimensional 53 tropical overturning circulation into its zonal and meridional components.

Keyser et al. [1989] have addressed the problem of partitioning three-dimensional cir-55 culations into orthogonal two-dimensional circulations. Their formalism is termed the 56 " $\psi$ -vector" method and has been applied to good effect in diagnosing the vertical circula-57 tions in mid-latitude baroclinic systems [Keyser et al., 1989; Reeder et al., 1991; Keyser 58 et al., 1992]. Loughe et al. [1995] have generalised the  $\psi$ -vector method to limited domains 59 and conformal map projections. In the present work, a version of the  $\psi$ -vector method 60 is used to partition the tropical atmosphere into Hadley and Walker circulations. With 61 the requirement that the circulations in the two orthogonal vertical planes satisfy conti-62

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<sup>63</sup> nuity independently, it is possible to uniquely attribute part of the vertical motion to the <sup>64</sup> circulation in the zonal direction and part to that in the meridional direction. This ap-<sup>65</sup> proach ensures that the complete three-dimensional circulation is equal to the sum of the <sup>66</sup> orthogonal two-dimensional circulations. The method will be used to define and examine <sup>67</sup> the *local Hadley* and *local Walker circulations*. Furthermore, it will be shown that the <sup>68</sup> approach is especially useful for defining regionally-averaged circulations.

The remainder of the paper is structured as follows. The method is outlined in Section 2. Section 3 discusses climatologies of the local Hadley and local Walker circulations. The anomalies that develop in association with El Niño and La Niña are presented in Section 4. This section examines also the relationship between these partitioned, orthogonal circulations and the more commonly used zonally and meridionally-averaged definitions of the Hadley and Walker circulations. Finally, the key conclusions of this study are summarised in Section 5.

## 2. Partitioning into zonal and meridional overturning circulations

This section describes the method by which the vertical motion, and hence, vertical 76 mass flux is partitioned into zonal and meridional overturning circulations. It will be 77 shown that the atmospheric vertical motion in pressure coordinates  $\omega$  can be written as 78 the sum of vertical motion associated with overturning in the zonal direction  $\omega_{\lambda}$ , and 79 vertical motion associated with overturning in the meridional direction  $\omega_{\phi}$ . The vertical 80 mass fluxes associated with these two components will be referred to here as the local 81 Walker circulation and the local Hadley circulation respectively, although the definitions 82 are different from those used by other authors. For example, Quan et al. [2004] define the 83 local circulations through an EOF analysis. 84

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## 2.1. A variation of the $\psi$ -vector method

<sup>85</sup> The  $\psi$ -vector method developed by *Keyser et al.* [1989] is an objective method to parti-<sup>86</sup> tion a three-dimensional irrotational flow into a unique pair of orthogonal, two-dimensional <sup>87</sup> circulations, with both satisfying continuity independently. *Keyser et al.* [1989], *Reeder* <sup>88</sup> *et al.* [1991] and *Keyser et al.* [1992] applied this method to fronts and cyclones on an f-<sup>89</sup> plane. In the present work, a variation of the  $\psi$ -vector method is reformulated in spherical <sup>80</sup> coordinates and used to partition globally the vertical mass flux into zonal and meridional <sup>81</sup> components (the local Walker and local Hadley circulations respectively).

The continuity equation in pressure coordinates is  $\nabla_p^2 \chi + \partial \omega / \partial p = 0$ , where p is the pressure,  $\nabla_p$  is the gradient operator along a pressure surface, and  $\chi$  is a velocity potential, which defines the divergent (irrotational) part of the horizontal wind through  $\mathbf{u}_{div} = \nabla_p \chi$ . A potential function  $\mu$  is defined such that  $\chi = \partial \mu / \partial p$ , which, when combined with the continuity equation, yields

$$\nabla_p^2 \mu = -\omega. \tag{1}$$

<sup>97</sup> This expression is identical to the potential function defined by *Eliassen* [1962] and <sup>98</sup> *Keyser et al.* [1989] (except the potential function  $\mu$  is denoted  $\chi$  in those studies). Equa-<sup>99</sup> tion (1) is a Poission equation which can be solved for  $\mu$  given suitable boundary condi-<sup>100</sup> tions.

Defining now the vector stream function  $\boldsymbol{\psi} = -\nabla_p \mu$  allows Eq. (1) to be written as

$$\nabla_p \cdot \boldsymbol{\psi} = \boldsymbol{\omega},\tag{2}$$

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where  $\boldsymbol{\psi}$  points towards (away from) regions of ascending (descending) motion. From the definitions of  $\mathbf{u}_{div}$ ,  $\chi$ , and  $\boldsymbol{\psi}$ , the divergent wind can be written in terms of the velocity streamfunction as

$$\mathbf{u}_{div} = (u_{\lambda}, u_{\phi}) = -\frac{\partial \psi}{\partial p} = -\left(\frac{\partial \psi_{\lambda}}{\partial p}, \frac{\partial \psi_{\phi}}{\partial p}\right),\tag{3}$$

where  $u_{\lambda}$  and  $u_{\phi}$  are the zonal and meridional components of the divergent part of the wind and  $\psi_{\lambda}$  and  $\psi_{\phi}$  are the components of  $\psi$  in the zonal and meridional directions respectively.

From Eqs. (1), (2) and (3) the vertical motion can be partitioned into the zonal and meridional planes as

$$\omega_{\lambda}\cos\phi = \frac{1}{a}\frac{\partial\psi_{\lambda}}{\partial\lambda} = \frac{1}{a^{2}\cos\phi}\frac{\partial^{2}\mu}{\partial\lambda^{2}},\tag{4}$$

$$\omega_{\phi}\cos\phi = \frac{1}{a}\frac{\partial}{\partial\phi}(\psi_{\phi}\cos\phi) = \frac{1}{a^2}\frac{\partial}{\partial\phi}\left(\cos\phi\frac{\partial\mu}{\partial\phi}\right),\tag{5}$$

where a is the radius of the Earth. Equations 3, 4 and 5 define two divergent circulations in orthogonal planes.

<sup>112</sup> Continuity is satisfied independently in both the zonal and meridional directions since

$$\frac{1}{a}\frac{\partial u_{\lambda}}{\partial \lambda} + \frac{\partial}{\partial p}(\omega_{\lambda}\cos\phi) = 0,$$

$$\frac{1}{a}\frac{\partial}{\partial\phi}(u_{\phi}\cos\phi) + \frac{\partial}{\partial p}(\omega_{\phi}\cos\phi) = 0$$

and consequently the vertical integral of the horizontal mass flux vanishes independently
in each orthogonal vertical plane. The vertical motion is the sum of the vertical motion

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partitioned in the two orthogonal directions  $\omega = \omega_{\lambda} + \omega_{\phi}$ , and the upward mass fluxes associated with the zonal and meridional parts of the circulation are

$$m_{\lambda} = -\omega_{\lambda} \cos\phi/g$$
 and  $m_{\phi} = -\omega_{\phi} \cos\phi/g$  (6)

respectively, where g is the gravitational acceleration.

Partitioning the three-dimensional overturning circulation into a pair of two-dimensional overturning circulations  $(u_{\lambda}, \omega_{\lambda})$  and  $(u_{\phi}, \omega_{\phi})$  is conceptually useful, objective, unambiguous, and especially beneficial if the circulations are averaged over a limited area as will be shown later. In the plots that follow the two-dimensional local Hadley circulation is portrayed by the vertical mass flux  $m_{\phi}$  and the divergent part of the circulation in the meridional plane  $(u_{\phi}, \omega_{\phi})$ . Likewise, the local Walker circulation is portrayed by the vertical mass flux  $m_{\lambda}$  and the divergent circulation in the zonal plane  $(u_{\lambda}, \omega_{\lambda})$ .

Note that the formulation outlined above is different from the  $\psi$ -vector method de-125 scribed by Keyser et al. [1989]. In the original formulation the horizontal wind is first 126 decomposed into its geostrophic and ageostrophic components  $\mathbf{u}_h = \mathbf{u}_g + \mathbf{u}_{ag}$ , and then 127 each component is subdivided into its rotational  $(\mathbf{u}_{g,rot}, \mathbf{u}_{ag,rot})$  and irrotational (divergent; 128  $\mathbf{u}_{g,irr}, \mathbf{u}_{ag,irr}$  parts. (Here the terms rotational and nondivergent are used interchangeably 129 as are the terms irrotational and divergent.) As the geostrophic wind is nondivergent, the 130 irrotational component of the geostrophic wind vanishes and the irrotational part of the 131 horizontal velocity is given only by the ageostrophic component. Hence, Keyser et al. 132 [1989] write the horizontal wind as  $\mathbf{u}_h = \mathbf{u}_{g,rot} + \mathbf{u}_{ag,rot} + \mathbf{u}_{ag,irr}$ . In the present work, the 133 horizontal wind is not separated into its geostrophic and ageostrophic components as the 134 method is applied to circulations in the tropics. The horizontal wind is separated only 135

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<sup>136</sup> into its rotational and irrotational components  $\mathbf{u}_h = \mathbf{u}_{irr} + \mathbf{u}_{rot}$ , which can be rewritten <sup>137</sup> as  $\mathbf{u}_h = -\nabla_p (\partial \mu / \partial p) + \mathbf{k} \times \nabla_p s$ , where  $(\partial \mu / \partial p)$  is the potential function,  $\mathbf{k}$  is the unit <sup>138</sup> vector in the vertical and s is a stream function. Helmholtz's theorem tells us that this <sup>139</sup> decomposition is unique on a sphere.

Indices closely related to the local zonal and local meridional overturning circulations 140 have been used previously in studies of the Asian monsoon Webster and Yang, 1992; Oort 141 and Yienger, 1996; Goswami et al., 1999] and the Hadley and Walker circulations [Moore 142 et al., 2004]. These indices, called the Zonal Overturning Index (ZOI) and Meridional 143 Overturning Index (MOI) by Moore et al. [2004], are defined by the difference in the 144 divergent part of the zonal and meridional winds at 200 hPa and 850 hPa, the expressions 145 for which are  $ZOI = u_{\lambda 200} - u_{\lambda 850}$  and  $MOI = u_{\phi 200} - u_{\phi 850}$ . The ZOI and MOI are 146 related to the components of the  $\psi$ -vector, but will not be used here. For reference, 147  $ZOI = \partial(\psi_{\lambda 200} - \psi_{\lambda 850})/\partial p$  and  $MOI = \partial(\psi_{\phi 200} - \psi_{\phi 850})/\partial p$  (using Eq. (3)). Replacing 148 the vertical derivative with a simple finite difference approximation, using Eqs. (4) and 149 (5), and assuming the vertical motion at 850 and 200 hPa is small compared with that 150 at 525 hPa (the center of the layer), gives  $m_{\lambda 525} = -\delta p/(2ag)\partial(ZOI)/\partial\lambda$  and  $m_{\phi 525} = -\delta p/(2ag)\partial(ZOI)/\partial\lambda$ 151  $-\delta p/(2ag\cos\phi)\partial(MOI\cos\phi)/\partial\phi$  where  $\delta p = 200 - 850$  hPa. With these assumptions, the 152 local Walker and local Hadley circulations are proportional to the zonal gradient of the 153 ZOI and the meridional gradient of the MOI respectively. 154

To conclude this section we note that several other studies have examined the zonal and meridional overturning circulations using the divergent part of the wind; these include Yuand Zwiers [2010], Hagos and Zhang [2010], and Hagos and Leung [2012]. However, only

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the present study shows how to decompose the divergent flow into two unique, orthogonal circulations that sum up to the original divergent flow filed.

## 2.2. Zonal and meridional averaging over restricted intervals

This section examines the relationship between the averaged overturning circulations and the local overturning circulations, which are part of the definitions of the Hadley and Walker circulations. Of course, averaging in meridional and zonal direction obscures the highly regional nature of the local Hadley and local Walker circulations.

The Hadley circulation (as distinct from the local Hadley circulation) is defined as the zonally-averaged circulation, where the average is taken around an entire circle of constant latitude. The zonally-averaged mass flux is  $[m]_{\lambda_1}^{\lambda_2} = [m_{\lambda}]_{\lambda_1}^{\lambda_2} + [m_{\phi}]_{\lambda_1}^{\lambda_2}$ , where the averaging operator is

$$[A]_{\lambda_1}^{\lambda_2} = \frac{1}{\Delta\lambda} \int_{\lambda_1}^{\lambda_2} A \, d\lambda. \tag{7}$$

Here, A is any given variable,  $\lambda_1$  and  $\lambda_2$  are the initial and final longitudes and  $\Delta \lambda = \lambda_2 - \lambda_1$ . In particular,  $[m_{\lambda}]_{\lambda_1}^{\lambda_2} = -(\psi_{\lambda}|_{\lambda_2} - \psi_{\lambda}|_{\lambda_1})/(ag\Delta\lambda\cos\phi)$ , a result which follows from the zonal average of Eqs. 4 and 5. In general,  $\psi_{\lambda}$  is small compared with  $\psi_{\phi}$  and, for this reason,  $[m_{\lambda}]_{\lambda_1}^{\lambda_2}$  is small compared with  $[m_{\phi}]_{\lambda_1}^{\lambda_2}$  even when the averaging is defined on a limited domain (see later Section 3.2).

That the relative smallness of  $[m_{\lambda}]_{\lambda_{1}}^{\lambda_{2}}$  is a consequence of the geometry of the sea surface temperature anomalies driving convection in the tropics can be illustrated by estimating the ratio between  $\omega_{\lambda}$  and  $\omega_{\phi}$  under the assumption that  $\cos \phi$  is almost one near the equator, in which case equations 4 and 5 give  $|\omega_{\lambda}/\omega_{\phi}| = |(\partial^{2}\mu/\partial\lambda^{2})/(\partial^{2}\mu/\partial\phi^{2})| \approx (\Delta\phi/\Delta\lambda)^{2}$ . Considering now the typical dimensions of the tongue of anomalously cool or anomalously

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<sup>178</sup> warm water in the eastern Pacific during ENSO events,  $\Delta \phi$  is of order 30 degrees and  $\Delta \lambda$ <sup>179</sup> is of order 100 degrees. In this case,  $|\omega_{\lambda}/\omega_{\phi}| \approx 0.1$ . This simple scale analysis shows that <sup>180</sup>  $\omega_{\lambda}$  is about one order of magnitude smaller than  $\omega_{\phi}$ .

<sup>181</sup> When averaging around an entire latitude circle we omit the terminals  $\lambda_2$  and  $\lambda_1$ , <sup>182</sup> simplifying the notation to [A]. As  $\lambda_2 = \lambda_1 + 2\pi$ ,  $[m_{\lambda}] = 0$ , and hence  $[m] = [m_{\phi}]$ , which <sup>183</sup> is consistent with previous definitions of the Hadley circulation. Of course,  $[m_{\lambda}]_{\lambda_1}^{\lambda_2}$  need not <sup>184</sup> vanish when the averaging operator is defined over a restricted domain. In this case, the <sup>185</sup> average mass flux  $[m]_{\lambda_1}^{\lambda_2}$  includes contributions from both  $\psi_{\lambda}$  and  $\psi_{\phi}$ , which complicates <sup>186</sup> the interpretation of the average circulation defined this way.

For reference, the commonly used mass-weighted meridional streamfunction [e.g. Hartmann, 1994] denoted here  $\Psi_{\phi}$ , is related to the meridional component of the  $\psi$ -vector through the expression  $\Psi_{\phi} = -(2\pi a \cos \phi/g)[\psi_{\phi}]$ , where the zonal average is taken from 0 to  $2\pi$ .

The Walker circulation (as distinct from the local Walker circulation) is defined as the meridionally-averaged circulation, although in this case it is common to restrict the averaging interval to the low latitudes. The meridional average of the mass flux is  $< m >_{\phi_1}^{\phi_2} = < m_\lambda >_{\phi_1}^{\phi_2} + < m_\phi >_{\phi_1}^{\phi_2}$ , where the averaging operator is now

$$\langle A \rangle_{\phi_1}^{\phi_2} = \frac{1}{\Delta \phi} \int_{\phi_1}^{\phi_2} A \cos \phi \, d\phi. \tag{8}$$

<sup>195</sup> Here  $\phi_1$  and  $\phi_2$  are the initial and final latitudes and  $\Delta \phi = \phi_2 - \phi_1$ . Specifically,  $< m_{\phi} >_{\phi_1}^{\phi_2} = -(\cos \phi \psi_{\phi}|_{\phi_2} - \cos \phi \psi_{\phi}|_{\phi_1})/(ag\Delta \phi)$ , which vanishes when  $\phi_2 = \pi/2$  and  $\phi_1 = -\pi/2$ , but is non-zero otherwise. It follows that when the Walker circulation is defined <sup>196</sup> over a restricted latitude band, the vertical mass flux includes a contribution from the

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<sup>199</sup> meridional circulation through the terms involving  $\psi_{\phi}$ . In general,  $\psi_{\phi}$  cannot be neglected <sup>200</sup> and, hence,  $\langle m_{\phi} \rangle_{\phi_1}^{\phi_2}$  makes a large, even dominant, contribution when averaging on a <sup>201</sup> limited domain (see Section 3.2).

For reference only, the relationship between the zonal component of the  $\psi$ -vector and the mass-weighted zonal streamfunction [e.g. *Hartmann*, 1994] is  $\Psi_{\lambda} = -(\pi a/g) < \psi_{\lambda} >$ , where  $\Psi_{\lambda}$  is the mass-weighted zonal streamfunction and the meridional average is taken from  $-\pi/2$  to  $\pi/2$ .

Henceforce, the meridional overturning circulation at a point or in a plane will be termed the *local Hadley circulation*, and the zonal overturning circulation at a point or in a plane will be termed the *local Waker circulation*. When the meridional overturning circulation and the zonal overturning circulations are averaged over an area, they will be referred to as *regional Hadley* and *regional Walker circulations*, respectively.

The regional Hadley (the zonally-averaged local Hadley circulation) and the regional Walker (meridionally-averaged local Walker circulation) circulations are defined as  $[m_{\phi}]_{\lambda_{1}}^{\lambda_{2}}$ and  $\langle m_{\lambda} \rangle_{\phi_{1}}^{\phi_{2}}$ , respectively, where  $(\lambda_{1}, \lambda_{2})$  and  $(\phi_{1}, \phi_{2})$  are arbitrary. The zonallyaveraged mass flux  $[m]_{\lambda_{1}}^{\lambda_{2}} = [m_{\phi}]_{\lambda_{1}}^{\lambda_{2}} + [m_{\lambda}]_{\lambda_{1}}^{\lambda_{2}}$  also includes a contribution from the zonal overturning circulation  $[m_{\lambda}]_{\lambda_{1}}^{\lambda_{2}}$ . Likewise, the contribution from the meridional overturning circulation to the meridionally-averaged mass flux  $\langle m \rangle_{\phi_{1}}^{\phi_{2}} = \langle m_{\phi} \rangle_{\phi_{1}}^{\phi_{2}} + \langle m_{\lambda} \rangle_{\phi_{1}}^{\phi_{2}}$ ,  $is \langle m_{\phi} \rangle_{\phi_{1}}^{\phi_{2}}$ .

## 3. Local Hadley and local Walker circulations

In this section the local Hadley circulation and the local Walker circulation are diagnosed from the ERA-Interim reanalysis [*Simmons et al.*, 2011; *Dee et al.*, 2011] with a horizontal resolution of 1.5 degrees and 37 vertical pressure levels for the period from 1979 to 2009.

## 3.1. Global climatology

The time-mean local Hadley and local Walker circulations as defined by the meridional and zonal mass fluxes at 500 hPa are shown in Figs. 1 and 2, respectively, for December, January, February (DJF) and June, July, August (JJA). The present study is not the first to characterize the Hadley and Walker circulations by the vertical motion [e.g. *Kumar et al.*, 1999; *Trenberth et al.*, 2000; *Wang*, 2002; *Tanaka et al.*, 2004], although it is the first to partition the vertical motion and attribute specific parts to the Hadley and Walker circulations.

In both seasons, the local Hadley circulation (Fig. 1) is characterized by zonally-228 elongated bands of vertical mass flux, with ascent prominent in the tropics and subsi-229 dence prominent in the subtropics, reflecting the mean rainfall (not shown). The ascent is 230 shifted towards the summer hemisphere and the subsidence is more marked in the winter 231 hemisphere. The main exception to this picture is in the equatorial eastern Pacific, where 232 there is subsidence along the equator, extending southwards into the extratropics. The 233 region is bounded to the north by a band of strong ascent marking the Inter-Tropical 234 Convergence Zone (ITCZ) there. 235

In contrast to the time-mean local Hadley circulation, the bands of vertical mass flux comprising the time-mean local Walker circulation are mostly oriented meridionally (Fig. 2). Moreover, almost everywhere, the time-mean local Walker circulation is weaker, and its seasonal variation smaller, than the corresponding time-mean local Hadley circulation. During Southern Hemisphere summer (DJF), the time-mean local Walker circulation (Fig. 2a) is most pronounced on the western sides on the Southern Hemisphere continents. The local Walker circulation is strongest in the eastern Pacific and South America, with

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<sup>243</sup> ascent over the continent and subsidence over the adjacent ocean. Similarly, there are <sup>244</sup> pronounced bands of subsidence over the Atlantic to the west of Africa, and over the <sup>245</sup> Indian Ocean to the West of Australia. These bands of subsidence are accompanied <sup>246</sup> by ascent of the western side of southern Africa and western Australia. A pattern of <sup>247</sup> alternating ascent and subsidence, most likely a consequence of the underlying orography, <sup>248</sup> is located over the Rockies, the Middle East and Asia.

The time-mean local Walker circulation in JJA (Fig. 2b) is more pronounced in the Northern Hemisphere. Centers of ascent are located in the Bay of Bengal and to its north, and over the south Pacific. Bands of ascent also lie along the eastern side of the north Pacific and through much of the Americas and the Caribbean. Notable centers of subsidence lie over the eastern Mediterranean Sea, over Somalia and the western Indian Ocean, to the east of the Caspian Sea, and off the west coast of the United States. Strong bands of ascent and subsidence lie along the Andes.

Vertical cross sections through the Hadley circulation  $[m_{\phi}]$  (zonally-averaged around 256 the whole globe) are shown in Fig. 3. Also plotted is the divergent circulation in the 257 plane of the cross section  $([u_{\phi}], [\omega_{\phi}])$ . In both seasons the Hadley circulation comprises an 258 ascending branch to the summer side of the equator accompanied by descending branches 259 to the north and south. The branch in the winter hemisphere is more pronounced than that 260 in the summer hemisphere. In JJA the downward mass flux in the Northern Hemisphere 261 is extremely weak effectively making the Hadley circulation single cell [James, 1994]. The 262 ascending branch of the Hadley circulation is wider and weaker during DJF and narrower 263 and more intense during JJA. As mentioned in Section 2.2, the contribution of the zonal 264

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overturning circulation  $[m_{\lambda}]$  vanishes when averaged zonally around the globe giving the textbook picture of the Hadley circulation [e.g. *Hartmann*, 1994; *James*, 1994].

## 3.2. Regional climatology

The principle advantage of identifying the local Hadley and local Walker circulations 267 through the  $\psi$ -vector method is that it leads to a simple unambiguous way to calculate 268 regional circulations. As an illustration, the regional Hadley circulation over the Maritime 269 Continent (zonally-averaged between 110°E and 160°E; box in Fig. 1)  $[m_{\phi}]_{110^{\circ}E}^{160^{\circ}E}$  and the 270 regional Walker circulation (meridionally-averaged between 10°N and 35°S; box in Fig. 271 2)  $< m_{\lambda} >_{35^{\circ}S}^{10^{\circ}N}$  are examined. The latter region is somewhat different from that typically 272 used to define the Walker circulation (e. g. 5°N to 5°S, which is sometimes referred to as 273 the equatorial Walker circulation [e.g. Bell and Halpert, 1998; Webster and Yang, 1992], 274 or 10°N-10°S [e.g. L'Heureux et al., 2013; Krishnamurthy and Goswami, 2000]). In the 275 following, both  $< m_{\lambda} >_{35^{\circ}S}^{10^{\circ}N}$  and  $< m_{\lambda} >_{5^{\circ}S}^{5^{\circ}N}$  are compared. The latitudinal interval from 276  $10^{\circ}$ N to  $35^{\circ}$ S is chosen because the maxima of the upward and downward zonal mass flux 277 lie in this band. 278

The regional Hadley circulation over the Maritime Continent  $[m_{\phi}]_{110^{\circ}E}^{160^{\circ}E}$  (Fig. 4) is differ-279 ent to the Hadley circulation  $[m_{\phi}]$  (Fig. 3). A clear double cell structure is found in DJF 280 (Fig. 4a), whereas in JJA (Fig. 4b) the single cell structure is apparent with the largest 281 downward mass flux mainly in the midlatitudes. In contrast to the Hadley circulation 282 (Fig. 3), the ascending branch of the regional Hadley circulation is broader in JJA than 283 during DJF. The descending branch in the winter hemisphere is more pronounced than in 284 the summer hemisphere. In JJA the descending branch comprises two maxima (see also 285 Fig. 1), the more southward associated with the descending branch of the regional Hadley 286

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circulation and the more northward is associated with the South Pacific convergence zone (SPCZ). These plots capture the annual shift in the position of the Hadley circulation and the influence of the Asian monsoon (JJA) and the Australian monsoon (DJF).

Figure 4c, d displays the contribution to the regional Hadley circulation from the zonal 290 overturning circulation  $[m_{\lambda}]_{110^{\circ}E}^{160^{\circ}E}$ . This contribution is the difference between the zonally-291 averaged mass flux  $[m]_{110^{\circ}E}^{160^{\circ}E}$  and the regionally-averaged Hadley circulation  $[m_{\phi}]_{110^{\circ}E}^{160^{\circ}E}$ , and 292 it represents the error in attributing the circulation to a regionally-averaged meridionally 293 overturning circulation. In DJF the zonal overturning circulation contributes to the as-294 cent south of about 20°N and to the descent north of this latitude. In JJA there is a 295 distinct contribution from the zonal overturning circulation to the upward vertical mass 296 flux. Overall, however, the contribution from the zonal overturning circulation to the 297 regional Hadley circulation is relatively small because the zonal overturning circulation 298 itself is small. In contrast, as discussed below, the meridional overturning circulation 299 contributes significantly to the regional Walker circulation, which is large compared to 300 the meridionally-averaged mass flux (see the scale analysis in section 2.2). 301

Figure 5 shows the regional Walker circulation defined by the latitudinal band [35°S, 302 10°N]. The three fields plotted are the regional Walker circulation  $\langle m_{\lambda} \rangle_{35^{\circ}S}^{10^{\circ}N}$ , the contri-303 bution to the circulation from the meridional overturning circulation  $\langle m_{\phi} \rangle_{35^{\circ}S}^{10^{\circ}N}$ , and the 304 sum  $\langle m_{\lambda} \rangle_{35^{\circ}S}^{10^{\circ}N} + \langle m_{\phi} \rangle_{35^{\circ}S}^{10^{\circ}N} = \langle m \rangle_{35^{\circ}S}^{10^{\circ}N}$ , which is the meridionally-averaged mass 305 flux. Throughout the year three distinct cells comprise the regional Walker circulation 306  $< m_{\lambda} >_{35^{\circ}S}^{10^{\circ}N}$ : one in the Indian Ocean, one in the Pacific Ocean and one in the Atlantic (see 307 the first row in Fig. 5). Also shown is the divergent circulation  $(\langle u_{\lambda} \rangle_{35^{\circ}S}^{10^{\circ}N}, \langle \omega_{\lambda} \rangle_{35^{\circ}S}^{10^{\circ}N})$ 308 attributable to the regional Walker circulation. The regional Walker circulation in the 309

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Pacific is the largest of the three, with ascent over the Maritime Continent and western 310 Pacific Ocean and descent in the eastern Pacific. In DJF the cell in the Pacific is very pro-311 nounced and the descent in the eastern Pacific Ocean is particularly strong. The regional 312 Walker circulation in the Atlantic is intense but much narrower, whereas the regional 313 Walker circulation in the Indian Ocean is relatively weak and most pronounced at the 314 west coast of Australia. In contrast, in JJA, the regional Walker circulation in the Indian 315 Ocean is much stronger and broader. The regional Walker circulation in the Atlantic is 316 weaker and the ascending branch of the Pacific cell is more distinct than the descending 317 branch. 318

The second row of Fig. 5 shows the meridionally-averaged mass flux  $< m >_{35^{\circ}S}^{10^{\circ}N}$  and 319 the wind field  $(\langle u_{\lambda} \rangle_{35^{\circ}S}^{10^{\circ}N}, \langle \omega \rangle_{35^{\circ}S}^{10^{\circ}N})$ . This is the circulation that results from simple 320 averaging. The differences between  $\langle m_{\lambda} \rangle_{35^{\circ}S}^{10^{\circ}N}$  and  $\langle m \rangle_{35^{\circ}S}^{10^{\circ}N}$  are shown in the third 321 row of Fig. 5. This difference quantifies the contribution from the meridional overturning 322 circulation  $\langle m_{\phi} \rangle_{35^{\circ}S}^{10^{\circ}N}$  and is the error in the regional Walker circulation when it is 323 identified with the mass flux  $< m >_{35^\circ S}^{10^\circ N}$ . In DJF there is a large positive contribution 324 from  $< m_{\phi} >_{35^{\circ}S}^{10^{\circ}N}$  to  $< m >_{35^{\circ}S}^{10^{\circ}N}$  over the Maritime Continent and the western Pacific as 325 well as over South America, which means that the values of  $\langle m_{\phi} \rangle_{35^{\circ}S}^{10^{\circ}N}$  reach up to 50% 326 of the values of  $< m >_{35^\circ S}^{10^\circ N}$ . Over the Maritime Continent, in particular,  $< m_{\phi} >_{35^\circ S}^{10^\circ N}$  is 327 larger than  $\langle m_{\lambda} \rangle_{35^{\circ}S}^{10^{\circ}N}$ . In JJA the contribution of  $\langle m_{\phi} \rangle_{35^{\circ}S}^{10^{\circ}N}$  to the meridionally-328 averaged downward vertical mass flux is large also. Here  $\langle m_{\phi} \rangle_{35^{\circ}S}^{10^{\circ}N}$  is up to one third 320 of  $< m >_{35^{\circ}S}^{10^{\circ}N}$ . Figure 5 underscores the point that partitioning the tropical circulation 330 into two orthogonal circulations is qualitatively simple but quantitatively less so. Simple 331 averaging over a restricted band of latitudes leads to a resulting meridionally-averaged 332

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mass flux  $< m >_{35^{\circ}S}^{10^{\circ}N}$  and circulation that are too strong, as part of the vertical mass flux should be attributed to the meridional overturning circulation, especially over the Maritime Continent and the Western Pacific. The reason for this is that the local Walker circulations are relatively weak compared to the local Hadley circulations.

It is common to define the regional Walker circulation as a meridional average between 337 5°S and 5°N [e.g. Bell and Halpert, 1998; Webster and Yang, 1992] (see Fig. 6). In 338 this case the regional Walker circulation  $\langle m_{\phi} \rangle_{5^{\circ}S}^{5^{\circ}N}$  (Fig. 6a, b) is slightly different to 339 that based on averaging between 35°S to 10°N  $< m_{\phi} >_{35^{\circ}S}^{10^{\circ}N}$  (Fig. 5a, b), although the 340 essential characteristics of the regional Walker circulations remain the same. However, 341 the differences between  $\langle m_{\lambda} \rangle_{5^{\circ}S}^{5^{\circ}N}$  and  $\langle m \rangle_{5^{\circ}S}^{5^{\circ}N}$  (Fig. 6c, d) are striking. Figure 6e, f 342 shows the contribution of the meridional overturning circulation  $\langle m_{\phi} \rangle_{5^{\circ}S}^{5^{\circ}N}$  to the zonally-343 averaged mass flux  $< m > 5^{\circ}N_{S}$ . The vertical mass flux is so strong partly because this 344 region lies in the ascending branch of the regional Hadley circulation, while the previous 345 region ( $35^{\circ}$ S and  $10^{\circ}$ N) was much broader and thus also contained parts of the descending 346 branches of the regional Hadley circulation leading to some cancelation. Nevertheless, the 347 contribution of the meridional overturning circulation in the band between  $5^{\circ}$ S and  $5^{\circ}$ N 348 is much stronger than in the band ranging from  $35^{\circ}$ S to  $10^{\circ}$ N. This indicates that, not 349 only is a careful partitioning of the vertical mass flux important, but the position of the 350 domain over which the average is taken affects the result. 351

## 4. ENSO and the local Hadley and local Walker circulations

The previous section showed that the partitioning of the vertical velocity and, hence, the vertical mass flux can be achieved in an objective and unambiguous way using the  $\psi$ -vector method. This method can be applied to a variety of problems, a particularly

interesting one being variations in the local Hadley and local Walker circulations with
 ENSO, the largest mode of variability in the Pacific region.

In this section anomalies from the local Hadley and local Walker circulation climatol-357 ogy are calculated for periods of El Niño and La Niña. To do so, first the time series 358 of the Southern Oscillation index (SOI), which was obtained from Australian Bureau 350 of Meteorology (ftp://ftp.bom.gov.au/anon/home/ncc/www/sco/soi/soiplaintext.html), 360 was smoothed using a three months running average. The smoothed SOI of  $\pm 0.9$  stan-361 dard deviation was used to separate the ERAI data into El Niño (47 months), La Niña 362 (51 months) and neutral (274 months) periods. Then seasonal averages were taken over 363 these periods. In the following we focuss on the El Niño and La Niña phases. 364

#### 4.1. Global ENSO climatology

Figures 7 and 8 show the mean anomalies at 500 hPa in the local Hadley and local Walker 365 circulations in DJF and JJA for El Niño and La Niña periods, respectively. During El 366 Niño events in DJF (Fig. 7a), the local Hadley circulation has a pronounced anomaly in 367 the ascent over the central Pacific, with strong anomalies in the subsidence to the north 368 and south of the ascent anomaly. There is a strong subsidence anomaly over the Maritime 369 Continent, weakening the mean ascent in that region (Fig. 1a). The subsidence anomaly 370 extends southeastward from the Maritime Continent across the South Pacific. In contrast, 371 during La Niña events (Fig. 7b), the ascent in the local Hadley circulation inceases over 372 the Maritime Continent and South Pacific (Fig. 1a). The response over the central Pacific 373 is both weaker and of the opposite sign to that during an El Niño. During El Niño 374 events in JJA (Fig. 7c) there is also a positive anomaly over the central Pacific. The 375 negative anomalies north and south of it are much weaker than in DJF. There is a band 376

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<sup>377</sup> of anomalous ascent across the Pacific islands. To the west of this band the subsidence <sup>378</sup> is enhanced. In JJA during La Niña, the signs of the anomalies are reversed and the <sup>379</sup> anomalies are generally weaker (Fig. 7d).

The main response of the local Walker circulation to an El Niño is a meridional band 380 of anomalous ascent in the central Pacific extending north and south along the eastern 381 margin of the basin (Fig. 8a). There are more confined bands of anomalous ascent running 382 from the Indian Ocean to the Middle East. A prominent band of anomalous subsidence 383 runs from the Maritime Continent both northwest to the western part of the north Pacific, 384 and southeast across Australia and the South Pacific. The response of the local Walker 385 circulation to a La Niña is less coherent than the response to an El Niño (Fig. 8b). There 386 is generally anomalous ascent across the Maritime Continent, as well as the regions north 387 and south of it, anomalous ascent in the Southwest Pacific, and anomalous subsidence 388 in the longitudes of India as well as in the Central Pacific. Thus, in agreement with 389 Oort and Yienger [1996], the local Walker circulation weakens during El Niño years and 390 strengthens during La Niña years. In JJA during an El Niño (Fig. 8c), the distribution of 391 the anomalies is very similar to that in DJF, although less coherent. There is an ascent 392 anomaly in the central Pacific and a descent anomaly along the coast of South America 393 and across the Maritime Continent. During a La Niña the picture of the local Walker 394 circulation in the Pacific is less clear (Fig. 8d). A band of anomalous descent lies along 395 the western Pacific in both hemispheres and a band of anomalous ascent in the Bay of 396 Bengal, over the Maritime Continent, and along the South American coastline. 397

Perhaps the key point to emerge from this analysis is that the local Hadley circulation responds more strongly to an El Niño than the local Walker circulation. The proba-

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ble reason for this difference in the response is that that region of warm sea water that 400 drives the El Niño is zonally elongated and, consequently, the temperature gradients in 401 the meridional direction are larger than those in the zonal direction. This interpretation 402 is consistent with the scale analysis given in Section 2.2. Hence, the  $\psi$ -vector approach 403 provides a viewpoint different from many introductory textbooks which portray ENSO 404 as a change in the Walker circulation. From this perspective, ENSO should not be por-405 trayed as a shift in the Walker circulation, but rather, a change in the Hadley circulation. 406 Nonetheless, in the central and eastern Pacific, the local Walker circulation is strengthened 407 during a La Niña and weakened during an El Niño. 408

#### 4.2. Regional Hadley circulation

The regional Hadley circulation response to ENSO over the Maritime Continent 400  $[m_{\phi}]_{110^{\circ}E}^{160^{\circ}E}$  is shown in Fig. 9a, b. In DJF during both El Niño and La Niña periods the 410 regional Hadley circulation has a typical two cell structure with ascent along the ITCZ 411 and descent on each side, with the winter cell being stronger. The ascending branch of 412 the regional Hadley circulation is stronger and wider during La Niña consistent with in-413 creased convection. The extent of the descending branches is the same in both periods. 414 The descending branch in the Northern Hemisphere shows marked differences. During El 415 Niño the descent is weaker and the maximum occurs at around 40°N. During La Niña, 416 this maximum in descent is stronger and a second maximum occurs at about 20°N and 417 at higher levels. 418

The zonally-averaged mass flux  $[m]_{110^{\circ}E}^{160^{\circ}E}$  shows a very similar structure (Fig. 9c, d) to the regional Hadley circulation. The differences between circulations are small, as shown in

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Fig. 9e, f which displays the contribution from the zonal overturning circulation  $[m_{\lambda}]_{110^{\circ}E}^{160^{\circ}E}$ . It is apparent that  $[m_{\lambda}]_{110^{\circ}E}^{160^{\circ}E}$  is much smaller that  $[m_{\phi}]_{110^{\circ}E}^{160^{\circ}E}$ .

In JJA, the regional Hadley circulation over the Maritime Continent  $[m_{\phi}]_{110^{\circ}E}^{160^{\circ}E}$  is char-423 acterized by a single cell structure. In contrast to DJF, the main ascent is located in the 424 same region during El Niño and La Niña periods, but is slightly stronger during El Niño 425 (Fig. 10a, b). The mid-latitude Ferrel cell in the summer hemisphere is due to averaging 426 on isobaric levels (see *Holton* [2004]). In the winter hemisphere during El Niño there are 427 two maxima in the descending branch of the regional Hadley circulation located close to 428 one another. During La Niña, however, the two maxima in descent are more distinct. The 429 equatorward maximum is related to the SPCZ whereas the poleward maximum is related 430 to the descending branch of the regional Hadley circulation. The contribution from the 431 zonal overturning circulation  $[m_{\lambda}]_{110^{\circ}E}^{160^{\circ}E}$  (Fig. 10e, f) to the zonally-averaged mass flux 432  $[m]_{110^{\circ}E}^{160^{\circ}E}$  (Fig. 10c, d) is mainly positive throughout the whole domain, with its maximum 433 in the region of ascent of the regional Hadley circulation. However, as in DJF, the con-434 tribution from the zonal overturning circulation is much smaller than the regional Hadley 435 circulation, which results in a close resemblance of  $[m_{\phi}]_{110^{\circ}E}^{160^{\circ}E}$  to  $[m]_{110^{\circ}E}^{160^{\circ}E}$ 436

## 4.3. Regional Walker circulations

Vertical cross-sections through the regional Walker circulations, defined by a meridional
average from 35°S to 10°N (as in Fig. 2), for DJF during El Niño and La Niña are shown
in Fig. 11a, b. During El Niño the strongest ascent occurs in the central Pacific with
the descending branch of the regional Walker circulation located in the eastern Pacific.
The regional Walker circulation in the Atlantic is much narrower although the ascent is
comparable in strength to that over the central Pacific. The maximum ascent occurs over

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the eastern part of South America which is accompanied by descent in the eastern Atlantic. 443 The regional Walker circulation in the Indian Ocean is the weakest. During La Niña, the 444 maximum ascent in the Pacific moves westward and is located over the western Pacific and 445 the Maritime Continent although the magnitude remains unchanged. The accompanying 446 descent in the eastern Pacific increases and the regional Walker circulation in the Pacific 447 becomes broader. The area of descent in the Indian Ocean extends westwards, and the 448 descent west of Australia is stronger. Moreover, the regional Walker circulation in the 440 Atlantic strengthens, but does not change position. 450

The Walker circulation based on the meridionally-averaged mass flux  $< m >_{35^{\circ}S}^{10^{\circ}N}$  is 451 much stronger than the regional Walker circulation  $\langle m_{\lambda} \rangle_{35^{\circ}S}^{10^{\circ}N}$  (Fig. 11c, d). The 452 contribution from the meridional overturning circulation  $\langle m_{\phi} \rangle_{35^{\circ}S}^{10^{\circ}N}$  is shown in Fig. 453 11e, f. It is particularly strong in the central Pacific and over South America during El 454 Niño and over the Maritime Continent and South America during La Niña. The Walker 455 circulation based on the divergent circulation has ascent over the Indian Ocean when the 456 regional Walker circulation has descent because of the effect of the meridional overturning 457 circulation. 458

The differences in the Walker circulation and the regional Walker circulation are even more dramatic when a merdional average between 5°S and 5°N is used (not shown). The meridionally-averaged mass flux  $< m > 5^{\circ} N_{5^{\circ} S}$  is too strong everywhere as the contribution from the meridional overturning circulation  $< m_{\phi} > 5^{\circ} N_{5^{\circ} S}$  is large, particularly in the regions of ascent. Note, however, that the meridional overturning circulation averaged between  $5^{\circ}$ S and  $5^{\circ}$ N,  $< m_{\lambda} > 5^{\circ} N_{5^{\circ} S}$ , and  $35^{\circ}$ S and  $10^{\circ}$ N,  $< m_{\lambda} > 10^{\circ} N_{5^{\circ} S}$ , provide very similar pictures.

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The regional Walker circulations are generally weaker in JJA than in DJF (not shown). 465 During El Niño the regional Walker circulation in the Indian Ocean is the strongest 466 whereas those in the Pacific region and in the Atlantic are relatively weak. During La Niña, 467 the region of ascent in the western Pacific and over the Maritime Continent remains in the 468 same location as in DJF but intensifies, leading to a stronger regional Walker circulation 469 in the Indian Ocean and in the Pacific. However, the regional Walker circulation in 470 the Indian Ocean is still the most distinct. As in DJF, the meridionally-averaged mass 471 flux produces a much stronger regional Walker circulation. The contribution from the 472 meridional overturning circulation is, however, weaker in JJA than in DJF, and larger in 473 the regions of descent than in the region of ascent. When averaging between 5°S and 5°N, 474 the contribution from the meridional overturning circulation is much stronger. 475

#### 5. Conclusions

A firmer, more precise and unambiguous basis for the common and useful practice of 476 partitioning of the three-dimensional tropical overturning circulation into the zonal and 477 meridional circulations has been sought as, in the past, the methods used to partition these 478 circulations have been mostly ad hoc. Using a simplified version of the  $\psi$ -vector method 479 developed by *Keyser et al.* [1989], the atmospheric vertical motion was partitioned into a 480 component associated with overturning in the zonal direction and a component associated 481 with overturning in the meridional direction. The mass fluxes associated with these two 482 components were called the local Walker circulation and the local Hadley circulation 483 respectively. This analysis is based on the divergent circulation as it is associated with 484 the vertical motion and hence, convection. The analysis focused on the period 1979 485 to 2009. Partitioning the three-dimensional overturning circulation into a pair of two-486

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dimensional overturning circulations proved to be conceptually useful and allowed for the decomposition of the divergent circulations in the zonal and meridional plane not only globally, but in specific regions. However, as the local Hadley and local Walker circulations were diagnosed from the ERA-Interim Reanalysis, the results must be treated with some caution as the accuracy of the divergent part of the circulation is uncertain, especially in the tropics.

The spatial pattern of the time-mean local Hadley circulation comprises localized, 493 zonally-elongated bands with mostly ascent in the tropics and mostly subsidence in the 494 subtropics, although in the eastern Pacific the mass flux is downward along the equa-495 tor and southwards into the extratropics. The regions of ascent are shifted towards the 496 summer hemisphere, while the regions of subsidence were more marked in the winter 497 hemisphere. The patterns of vertical mass flux comprising the time-mean local Walker 498 circulation are mostly oriented meridionally. Almost everywhere, the time-mean local 499 Hadley circulation is stronger, and its seasonal variation larger, than the corresponding 500 time-mean local Walker circulation. During DJF, the time-mean local Walker circula-501 tion is strongest in the eastern Pacific and South America, with upward mass fluxes over 502 the continent and downward mass fluxes over the adjacent ocean. The time-mean local 503 Walker circulation is prominent also through Asia and eastern Africa during JJA. 504

The local Walker circulation is stronger during a La Niña and weaker during an El Niño, especially in the eastern Pacific. However, the local Hadley circulation responds more strongly to an El Niño than the local Walker circulation, suggesting that an El Niño cannot be interpreted as simply a zonal shift in the Walker circulation.

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The Hadley and Walker circulations (as distinct from the local Hadley and local Walker 509 circulations) are commonly defined as the zonal and meridional averages of the overturning 510 circulation respectively. In defining the Walker circulation, the average is often defined 511 over a restricted latitudinal band. It was shown that, in this case, the vertical mass 512 flux includes a contribution from the meridional overturning circulation. Nonetheless, 513 defining the Hadley circulation by averaging over a limited region introduces only small 514 errors as the contribution from the zonal overturning circulation is small. In contrast, 515 the Walker circulation should not generally be defined by averaging over a limited re-516 gion, as the meridional overturning circulation contributes to the definition of the Walker 517 circulation. To avoid this problem the regional Hadley and regional Walker circulations 518 were defined as the meridional and zonal averages of the local Hadley and local Walker 519 circulations respectively. The contribution from the meridional overturning circulation to 520 the meridionally-averaged mass flux becomes even more acute when averaging between 521 5°S and 5°N instead of 35°S and 10°N. This emphasizes the dependance of the result on 522 the partitioning of the vertical mass flux as well as the region of averaging. 523

As ENSO is associated with the largest interannual variability in the Pacific region we 524 investigated the regional Hadley circulation over the Maritime Continent as well as the 525 regional Walker circulations in each ocean basin during El Niño and La Niña phases. 526 In DJF, the regional Hadley circulation over the Maritime Continent is stronger and the 527 region of ascent is wider during La Niña than during El Niño. In JJA, the regional Hadley 528 circulation is stronger during El Niño, although the cell in the Northern Hemisphere is 529 weaker during El Niño than during La Niña. The cell of the regional Hadley circulation 530 located in the Northern Hemisphere is relatively weak during both El Niño and La Niña. 531

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In addition there is a double cell structure for the regional Hadley circulation in DJF and a single cell structure in JJA.

In DJF during an El Niño, the regional Walker circulation in the Pacific is characterized 534 by the region of strongest ascent in the central Pacific. During a La Niña the ascent is 535 shifted westwards although the magnitude remains unchanged. The descent is located 536 over the eastern Pacific and becomes stronger during La Niña. Thus, the regional Walker 537 circulation in the Pacific becomes wider. During El Niño the regional Walker circulation 538 in the Pacific is relatively narrow, although it is still the largest regional Walker circu-539 lation. In the Indian Ocean, the regional Walker circulation is relatively weak with the 540 largest descent west of the western Australian coast during La Niña. The regional Walker 541 circulation in the Atlantic is narrow, but of similar intensity to that in the Pacific, and 542 intensifies during La Niña. In JJA, the regional Walker circulation in the Indian Ocean 543 is strongest, particularly during La Niña. The regional Walker circulation in the Pacific 544 is stronger during La Niña but remains in roughly the same place with the main ascent 545 in the western Pacific.

As a next step we will use the  $\psi$ -vector method to analyze trends in the regional Hadley and regional Walker circulations. We will also apply the  $\psi$ -vector method to decompose the quasi-geostrophic omega equation and to attribute the source terms of this equation to changes in the regional Hadley and regional Walker circulations.

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Acknowledgments. This research was supported by the Australian Research Council Grant FS100100081. We would like to thank Roger Smith for his valuable comments on an earlier version of this manuscript.

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**Figure 1.** The local Hadley circulation. The meridional mass flux  $(kg m^{-2} s^{-1})$  at 500 hPa calculated from the ERAI reanalysis (1979 - 2009) for the seasons (a) DJF and (b) JJA is shown in shadings of red (ascent) and blue (descent). The black lines at 110°E and 160°E mark the area defining the regional Hadley circulation over the Maritime Continent.



Figure 2. The local Walker circulation. The zonal mass flux  $(kg m^{-2} s^{-1})$  at 500 hPa calculated from the ERAI reanalysis (1979 - 2009) for the seasons (a) DJF and (b) JJA. The box shown (0-360°, 10°N-35°S) marks the area defining the regional Walker circulations.

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Figure 3. The Hadley circulation from the ERAI reanalysis (1979 - 2009) for (a) DJF and (b) JJA. The vertical mass flux in meridional direction  $([m_{\phi}], \text{kg m}^{-2} \text{s}^{-1})$  is shaded. Vectors represent the wind in the plane of the cross section  $([u_{\phi}], [\omega_{\phi}])$ . The zero value of  $[m_{\phi}]$  is displayed as a thick black line. Fields below the mean orography are omitted.



Figure 4. The regional Hadley circulation over the Maritime Continent (region shown in Fig. 1) from the ERAI reanalysis (1979 - 2009) for (a) DJF and (b) JJA averaged over the longitudes 110 °E to 160 °E. The vertical mass flux in the meridional plane  $([m_{\phi}]_{110^{\circ}E}^{160^{\circ}E}, \text{ kg m}^{-2} \text{ s}^{-1})$  is shaded. Vectors represent the wind in the plane of the cross section  $([u_{\phi}]_{110^{\circ}E}^{160^{\circ}E}, [\omega_{\phi}]_{110^{\circ}E}^{160^{\circ}E})$ . The contribution to the vertical mass flux from the zonal overturning circulation  $([m_{\lambda}]_{110^{\circ}E}^{160^{\circ}E} = [m]_{110^{\circ}E}^{160^{\circ}E} - [m_{\phi}]_{110^{\circ}E}^{160^{\circ}E}, \text{ kg m}^{-2} \text{ s}^{-1})$  for (c) DJF and (d) JJA is shaded. The zero value of the respective variable is displayed as a thick black line. Fields below the maximum orography are omitted.

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Figure 5. The regional Walker circulation (region shown in Fig. 2) from the ERAI reanalysis (1979 - 2009) for (a) DJF and (b) JJA averaged over the latitudes 35 °S to 10 °N. The vertical mass flux in the zonal plane ( $< m_{\lambda} > ^{10^{\circ}N}_{35^{\circ}S}$ , kg m<sup>-2</sup> s<sup>-1</sup>) is shaded. Vectors represent the wind in the plane of the cross section ( $< u_{\lambda} > ^{10^{\circ}N}_{35^{\circ}S}$ ,  $< \omega_{\lambda} > ^{10^{\circ}N}_{35^{\circ}S}$ ). The vertical mass flux ( $< m > ^{10^{\circ}N}_{35^{\circ}S}$ , kg m<sup>-2</sup> s<sup>-1</sup>) is shaded for (c) DJF and (d) JJA. Vectors represent the wind in the plane of the cross section ( $< u_{\lambda} > ^{10^{\circ}N}_{35^{\circ}S}$ ,  $< \omega > ^{10^{\circ}N}_{35^{\circ}S}$ ). The contribution to the vertical mass flux from the meridional overturning circulation  $< m_{\phi} > ^{10^{\circ}N}_{35^{\circ}S} = < m > ^{10^{\circ}N}_{35^{\circ}S} - < m_{\lambda} > ^{10^{\circ}N}_{35^{\circ}S}$  (kg m<sup>-2</sup> s<sup>-1</sup>) is shaded for (e) DJF and (f) JJA. The zero value of the respective fields is displayed as a thick black line. Fields below the A F T maximum orography are omitted.



Figure 6. The same as Fig. 5, but meridionally averaged between 5°S and 5°N.



Figure 7. Anomalies of the local Hadley circulations at 500 hPa with respect to the mean for 1979-2009. The anomalies in the meridional mass flux  $m_{\phi}$  (kg m<sup>-2</sup> s<sup>-1</sup>) for El Niño (left) and La Niña (right) periods are shown for DJF (top row) and JJA (bottom row).



Figure 8. Anomalies of the local Walker circulation at 500 hPa with respect to the mean for 1979-2009. The anomalies in the zonal mass flux  $m_{\lambda}$  (kg m<sup>-2</sup> s<sup>-1</sup>) for El Niño (left) and La Niña (right) periods are shown for DJF (top row) and JJA (bottom row).



**Figure 9.** The regional Hadley circulation for the Maritime Continent (see box in Fig. 1) during El Niño (left) and La Niña (right) periods in DJF calculated from the ERAI reanalysis (1979 - 2009) averaged over the longitudes 110 °E to 160 °E. (a), (b): The regional Hadley circulation  $([m_{\phi}]_{110^{\circ}E}^{160^{\circ}E}, \text{kg m}^{-2} \text{s}^{-1})$  is shaded, and the wind in the plane of the cross section  $([u_{\phi}]_{110^{\circ}E}^{160^{\circ}E}, [\omega_{\phi}]_{110^{\circ}E}^{160^{\circ}E})$ . (c), (d): The vertical mass flux  $([m]_{110^{\circ}E}^{160^{\circ}E}, \text{kg m}^{-2} \text{s}^{-1})$  is shaded, and the wind in the plane of the cross section  $([u_{\phi}]_{110^{\circ}E}^{160^{\circ}E}, [\omega_{\phi}]_{110^{\circ}E}^{160^{\circ}E})$ . (e), (f): The contribution of the zonal overturning circulation  $([m_{\lambda}]_{110^{\circ}E}^{160^{\circ}E}, \text{kg m}^{-2} \text{s}^{-1})$  to the vertical mass flux is shaded. In all plots the zero-line is displayed as a thick black line. The maximum orography in this box is masked out. D R A F T December 12, 2013, 1:46pm D R A F T



Figure 10. The same as Fig. 9, but for JJA.



**Figure 11.** The regional Walker circulation for El Niño (left) and La Niña (right) periods in DJF calculated from the ERAI reanalysis (1979 - 2009) averaged over the latitudes 35 °S to 10 °N. (a), (b): The vertical mass flux in the zonal plane  $(< m_{\lambda} >_{35^{\circ}S}^{10^{\circ}N}, \text{kg m}^{-2} \text{s}^{-1})$  is shaded, and the wind in the plane of the cross section  $(< u_{\lambda} >_{35^{\circ}S}^{10^{\circ}N}, < \omega_{\lambda} >_{35^{\circ}S}^{10^{\circ}N})$ . (c), (d): The vertical mass flux  $(< m >_{35^{\circ}S}^{10^{\circ}N}, \text{kg m}^{-2} \text{s}^{-1})$  is shaded, and the wind in the plane of the cross section  $(< u_{\lambda} >_{35^{\circ}S}^{10^{\circ}N}, < \omega >_{35^{\circ}S}^{10^{\circ}N})$ . Note, the magnitude of the wind vectors is double the magnitude of the wind vectors shown in (a) and (b). (e), (f): The contribution from the meridional circulation  $(< m_{\phi} >_{35^{\circ}S}^{10^{\circ}N}, \text{kg m}^{-2} \text{s}^{-1})$  to the mass flux is shaded. In all plots the zero-line is displayed as a thick black line. The maximum orography in this D R A F T December 12, 2013, 1:46pm D R A F T box is masked out.



Figure 12. The same as Fig. 11, but for JJA, except the wind vectors all have the same reference vector.